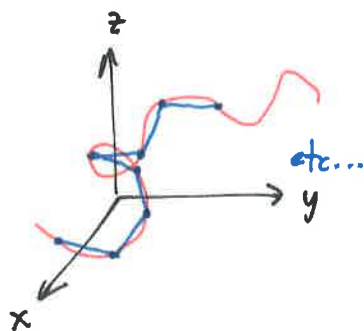


## Geometry of Curves in Space : 10.8 & 10.9



$\gamma$  or  $C$  : space curve - graph of vector function.

Suppose the curve is parametrized:

$$\vec{r}(t) = c(t) = \langle x(t), y(t), z(t) \rangle$$

since curves and vector functions are so closely related, we confuse them.

Recall from Chapters 7/9: the arc length of a 2D-curve is given by:

$$L(c| [a, b]) = \int_a^b \sqrt{(\dot{x})^2 + (\dot{y})^2} dt = \int_a^b \sqrt{\left(\frac{dx}{dt}\right)^2 + \left(\frac{dy}{dt}\right)^2} dt$$

In 3D, the formula is the same:

$$L(c| [a, b]) = \int_a^b \sqrt{(\dot{x})^2 + (\dot{y})^2 + (\dot{z})^2} dt = \int_a^b \|\dot{c}(t)\| dt$$

Ex.  $c(t) = \cos t \vec{i} + \sin t \vec{j} + t \vec{k}$

Find the arc length from  $(1, 0, 0)$  to  $(1, 0, 2\pi)$

i.e.,  $t : 0 \rightarrow 2\pi$

~~$$\|\dot{c}(t)\| = \sqrt{\cos^2 t + \sin^2 t + 1}$$~~

$$\dot{c}(t) = -\sin t \vec{i} + \cos t \vec{j} + 1 \vec{k}$$

$$\|\dot{c}(t)\| = \sqrt{\sin^2 t + \cos^2 t + 1} = \sqrt{2}$$

$$L(c| [a, b]) = \int_0^{2\pi} \sqrt{2} dt = \boxed{2\pi\sqrt{2}} \quad \text{u}$$

Recall that a single curve  $c$  can be parametrized in many different ways. Some are more useful than others.

Also recall the arc length function:

$$s(t) = \int_a^t \|\dot{c}(u)\| du \quad (\text{via the FTC})$$



$s$  gives the arc length from a set point  $c(a)$  to the point  $c(u)$ .

Also by the FTC:  $\frac{d}{dt} s(t) = \frac{d}{dt} \int_a^t \|\dot{c}(u)\| du = \|\dot{c}(t)\|$

or  $\dot{s}(t) = \|\dot{c}(t)\|$ .

It's often useful to parametrize by arc length.

$s(t)$  is an increasing one-to-one function, so we can solve for  $t$  as a function of  $s$ :  $t = t(s)$ . Then do a change of variables to write  $c = c(t(s))$ .

Then, if  $s=3$   $c(t(3)) = \vec{r}(t(3))$  is the position vector of the point 3 units of arc length away from the starting point of the curve.

Ex. Reparametrize the helix  $c(t) = \langle \cos t, \sin t, t \rangle$  by arclength.  
w/ initial point  $(1, 0, 0)$ , in the direction of increasing  $t$ .

From last Ex,  $\dot{c}(t) = \|\dot{c}(t)\| = \sqrt{2}$

$$\text{so } s(t) = \int_0^t \sqrt{2} \, du = \sqrt{2} t$$

$$\text{then } t = \frac{s}{\sqrt{2}}$$

The reparametrized curve is then

$$c(\frac{s}{\sqrt{2}}) = \left\langle \cos\left(\frac{s}{\sqrt{2}}\right), \sin\left(\frac{s}{\sqrt{2}}\right), \frac{s}{\sqrt{2}} \right\rangle. \quad \text{Word.}$$

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### Curvature:

Recall the unit tangent vector field on a curve  $c(t) = \vec{r}(t)$ :

$$T(t) = \frac{\dot{c}(t)}{\|\dot{c}(t)\|}$$

indicates the direction of curve at that  $t$ -value.

Notice:  $T$  changes very slowly if  $c$  is nearly straight.

$T$  changes more quickly if  $c$  is very "curvy".

Thus, we define the curvature of  $c$  as

$$\kappa = \left\| \frac{dT}{ds} \right\| = \frac{\|\ddot{c}\|}{\|\dot{c}\|^3} (t) = \kappa(t).$$

Ex. Show that the curvature of a circle w/ radius  $a$  is  $\frac{1}{a}$ .

$$c(t) = \langle a \cos t, a \sin t \rangle$$

$$\dot{c}(t) = \langle -a \sin t, a \cos t \rangle$$

$$\|\dot{c}(t)\| = \sqrt{a^2 \sin^2 t + a^2 \cos^2 t} = |a| = a$$

$$T(t) = \frac{\dot{c}}{\|\dot{c}\|} = \langle -\sin t, \cos t \rangle$$

$$\text{Thus, } \kappa(t) = \frac{\|\dot{T}\|}{\|\dot{c}\|} = \frac{1}{a}.$$

$$\dot{T} = \langle -\cos t, -\sin t \rangle$$

$$\|\dot{T}\| = \sqrt{\cos^2 t + \sin^2 t} = 1.$$

Theorem: The curvature of a curve  $c$  w/ pos. function  $\hat{r}$  is

$$\kappa(t) = \frac{\|\dot{\hat{r}}(t) \times \ddot{\hat{r}}(t)\|}{\|\dot{\hat{r}}(t)\|^3} = \frac{\|\dot{c}(t) \times \ddot{c}(t)\|}{\|\dot{c}(t)\|^3}$$

Proof: RE. Follow along w/ work in book. p. 573.  $\square$

Ex.  $\vec{r}(t) = \langle t, t^2, t^3 \rangle$  Find  $\kappa$  at  $\langle 0, 0, 0 \rangle$ .

$$\dot{\vec{r}}(t) = \langle 1, 2t, 3t^2 \rangle$$

$$\dot{\vec{r}} \times \ddot{\vec{r}} = \langle 12t^2 - 6t^2, 0 - 6t, 2 \rangle = \langle 6t^2, -6t, 2 \rangle$$

$$\ddot{\vec{r}}(t) = \langle 0, 2, 6t \rangle$$

$$\|\dot{\vec{r}} \times \ddot{\vec{r}}\| = \sqrt{36t^4 + 36t^2 + 4} \Rightarrow \kappa(0) = \frac{\sqrt{4}}{1^{3/2}} = 2.$$

$$\|\dot{\vec{r}}\|^3 = (1 + 4t^2 + 9t^4)^{3/2}$$

## Normal and binormal vectors:

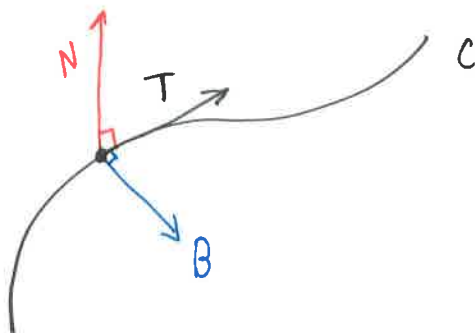
Recall  $\|T(t)\| = 1$  for all  $t$ , and  $T(t) \cdot \dot{T}(t) = 0$ . (Ex from last class.)

This means  $\dot{T} \perp T$  for all  $t$ .

$\dot{T}$  is not necessarily a unit vector itself, but

$$N(t) = \frac{\dot{T}(t)}{\|\dot{T}(t)\|} \text{ is.}$$

$N(t)$  is the principal unit normal vector of  $c = \vec{r}(t)$ .



Two vectors determine a plane, which then has a normal vector, the binormal vector to the curve:

$$B(t) := T(t) \times N(t)$$

Together  $\{T, N, B\}$  form a frame or basis for  $\mathbb{R}^3$  centered at the point on the curve  $c$ .

Defn. The plane determined by  $N$  and  $B$  is called the normal plane to  $C$  at  $P$ . It contains all vectors orthogonal to  $C$ .

The plane determined by  $T$  and  $N$  is called the osculating plane of  $C$  at  $P$ . It "kisses" the curve at  $P$ . The plane is "tangent" to the curve.

Ex. Find the normal and binormal vectors for the helix:

$$\vec{r}(t) = \langle \cos t, \sin t, t \rangle$$

$$\dot{\vec{r}}(t) = \langle -\sin t, \cos t, 1 \rangle$$

$$\|\dot{\vec{r}}(t)\| = \sqrt{2}$$

$$T(t) = \frac{\dot{\vec{r}}(t)}{\|\dot{\vec{r}}(t)\|} = \frac{1}{\sqrt{2}} \langle -\sin t, \cos t, 1 \rangle$$

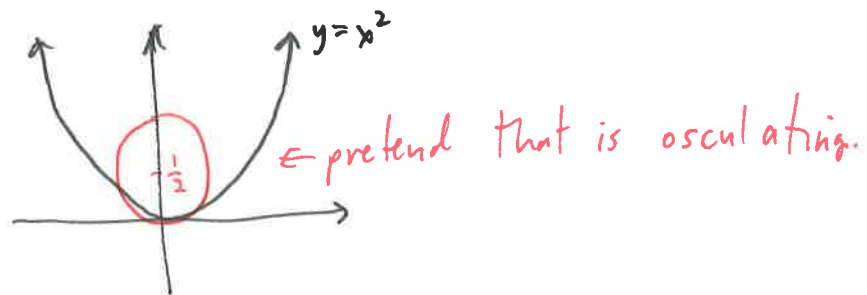
$$\ddot{\vec{r}}(t) = \frac{1}{\sqrt{2}} \langle -\cos t, -\sin t, 0 \rangle$$

$$\|\ddot{\vec{r}}(t)\| = \frac{1}{\sqrt{2}} \sqrt{\cos^2 t + \sin^2 t} = \frac{1}{\sqrt{2}}$$

$$\text{so, } N(t) = \frac{\ddot{\vec{r}}(t)}{\|\ddot{\vec{r}}(t)\|} = \langle -\cos t, -\sin t, 0 \rangle$$

$$B(t) = T(t) \times N(t) = \frac{1}{\sqrt{2}} \begin{vmatrix} \vec{i} & \vec{j} & \vec{k} \\ -\sin t & \cos t & 1 \\ -\cos t & -\sin t & 0 \end{vmatrix} = \frac{1}{\sqrt{2}} \langle \sin t, -\cos t, 1 \rangle = B$$

Another Defn. The circle that lies in the osculating plane, ~~and~~ has the same tangent to  $c$  at  $P$ , lies on the concave side of  $c$  (toward  $N$ ) and has radius  $\frac{1}{\kappa} = \rho$  is called the osculating circle of  $c$  at  $P$ .




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Curves as paths of particles: Motion in space

$\vec{r}(t) = \langle x(t), y(t), z(t) \rangle$  is the position function of a particle moving in space.

The velocity is then

$$\vec{v}(t) := \dot{\vec{r}}(t) = \langle \dot{x}(t), \dot{y}(t), \dot{z}(t) \rangle,$$

speed is

$$\|\vec{v}(t)\| = \|\dot{\vec{r}}(t)\| = \frac{ds}{dt},$$

and acceleration is

~~$\vec{a}(t) := \|\dot{\vec{v}}(t)\| = \|\ddot{\vec{r}}(t)\|$~~   
whoops!!

$$\vec{a}(t) := \dot{\vec{v}}(t) = \ddot{\vec{r}}(t).$$

Ex.  $\vec{r}(t) = \langle t^3, t^2 \rangle$  Find  $\vec{v}$ ,  $\|\vec{v}\|$ ,  $\vec{a}$  when  $t=1$ .

$$\left\{ \begin{array}{l} \vec{v} = \dot{\vec{r}} = \langle 3t^2, 2t \rangle \\ \vec{v}(1) = \langle 3, 2 \rangle \\ \|\vec{v}\| = \sqrt{9+4} = \sqrt{13} \\ \vec{a} = \dot{\vec{v}} = \langle 6t, 2 \rangle \\ \vec{a}(1) = \langle 6, 2 \rangle \end{array} \right.$$

Ex.  $\vec{a} = \langle 4t, 6t, 1 \rangle$  Find  $\vec{r}(t)$ .

$$\vec{v}(0) = \langle 1, -1, 1 \rangle$$

$$\vec{r}(0) = \langle 1, 0, 0 \rangle$$

$$\begin{aligned} \vec{v}(t) &= \int \vec{a}(t) dt = \int \langle 4t, 6t, 1 \rangle dt \\ &= \langle 2t^2, 3t^2, t \rangle + \vec{c} \end{aligned}$$

$$\vec{v}(0) = \langle 0, 0, 0 \rangle + \vec{c} \Rightarrow \vec{c} = \langle 1, -1, 1 \rangle$$

$$\text{so } \vec{v}(t) = \langle 2t^2+1, 3t^2-1, t+1 \rangle$$

$$\begin{aligned} \vec{r} &= \int \vec{v} dt = \int \langle 2t^2+1, 3t^2-1, t+1 \rangle dt \\ &= \left\langle \frac{2}{3}t^3+t, t^3-t, \frac{1}{2}t^2+t \right\rangle + \vec{d} \end{aligned}$$

$$\vec{r}(0) = \langle 1, 0, 0 \rangle = \langle 0, 0, 0 \rangle + \vec{d} \Rightarrow \vec{d} = \langle 1, 0, 0 \rangle$$

$$\text{so } \boxed{\vec{r}(t) = \left\langle \frac{2}{3}t^3+t+1, t^3-t, \frac{1}{2}t^2+t \right\rangle}$$

Newton's Second Law of motion:

$$\vec{F}(t) = m \vec{a}(t) \quad \text{where } m \text{ is the mass of the particle.}$$

HW may use this.

Tangential and normal acceleration:

let speed be  $\sigma(t) := \|\vec{v}(t)\|$  (book uses  $v$ , but that's confusing.)

$$\vec{T}(t) = \frac{\dot{\vec{r}}(t)}{\|\dot{\vec{r}}(t)\|} = \frac{\vec{v}(t)}{\|\vec{v}(t)\|} = \frac{\vec{v}(t)}{\sigma(t)} \quad \text{or, just } \frac{\vec{v}}{\sigma}$$

$$\text{Thus } \underline{\vec{v} = \sigma \vec{T}} \Rightarrow \underline{\vec{a} = \dot{\vec{v}}} = \sigma' \vec{T} + \sigma \dot{\vec{T}} \quad (*)$$

$$\text{Recall } \kappa = \frac{\|\dot{\vec{T}}\|}{\|\vec{v}\|} = \frac{\|\dot{\vec{T}}\|}{\sigma}, \quad \text{so } \|\dot{\vec{T}}\| = \kappa \sigma$$

$$\text{Also, } \vec{N} = \frac{\dot{\vec{T}}}{\|\dot{\vec{T}}\|}, \quad \text{so } \dot{\vec{T}} = \|\dot{\vec{T}}\| \vec{N} = \kappa \sigma \vec{N}$$

Then (\*) becomes

$$\vec{a}(t) = \underbrace{\sigma' \vec{T}}_{\text{tangential}} + \underbrace{\kappa \sigma^2 \vec{N}}_{\text{normal}}$$

For simplicity we write:  $\vec{a} = a_T \vec{T} + a_N \vec{N}$ .

It turns out (see book pp. 582f)

$$a_T = \sigma' = \frac{\dot{\vec{r}} \cdot \ddot{\vec{r}}}{\|\dot{\vec{r}}\|}$$

and

Fill in details, if time.

$$a_N = \kappa r^2 = \frac{\|\dot{\vec{r}} \times \ddot{\vec{r}}\|}{\|\dot{\vec{r}}\|^2}$$

Ex. A particle moves w/ position function  $\vec{r}(t) = \langle t^2, t^2, t^3 \rangle$

Find  $a_T$  and  $a_N$ .

$$a_T = \frac{\dot{\vec{r}} \cdot \ddot{\vec{r}}}{\|\dot{\vec{r}}\|}$$

$$\dot{\vec{r}} = \langle 2t, 2t, 3t^2 \rangle$$

$$\ddot{\vec{r}} = \langle 2, 2, 6t \rangle$$

$$\dot{\vec{r}} \cdot \ddot{\vec{r}} = 4t + 4t + 18t^3 = 8t + 18t^3$$

$$\Rightarrow a_T = \frac{8t + 18t^3}{\sqrt{8t^2 + 9t^4}}$$

$$\|\dot{\vec{r}}\| = \sqrt{4t^2 + 4t^2 + 9t^4} = \sqrt{8t^2 + 9t^4}$$

$$a_N = \frac{\|\dot{\vec{r}} \times \ddot{\vec{r}}\|}{\|\dot{\vec{r}}\|^2}$$

$$\dot{\vec{r}} \times \ddot{\vec{r}} = \langle 12t^2 - 6t^2, 6t^2 - 12t^2, 0 \rangle = \langle 6t^2, -6t^2, 0 \rangle$$

$$\|\dot{\vec{r}} \times \ddot{\vec{r}}\| = \sqrt{36t^4 + 36t^4} = \sqrt{72t^4} = \sqrt{72} t^2$$

$$\Rightarrow a_N = \frac{\sqrt{72} t^2}{\sqrt{8t^2 + 9t^4}}$$

## Chapter 10: Grand Finale

### Kepler's Laws of Planetary Motion:

1. A planet revolves around the sun in an elliptical orbit w/ the sun at one focus.
2. The line joining the sun to a planet sweeps out equal areas in equal times.
3. The square of the period of revolution of a planet is proportional to the cube of the length of the major axis of its orbit.

Extra Credit Project: Kepler project from book. Up to 2% on Final Grade. Due by last day of class (9 May). Will be graded slightly harder than normal projects.